### **Neutrino Self-Interactions**

Nikita Blinov

September 17, 2020

NF03 Kick Off Meeting





Based on Lol with Mauricio Bustamante, Kevin Kelly and Yue Zhang SNOWMASS21-NF3-003

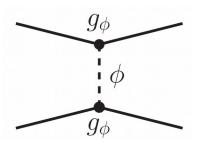
### **Self-Interactions in the SM**

 Neutrino self-interactions are some of the most difficult operators to probe in SM EFT

$$\frac{G_F\rho}{\sqrt{2}}(\bar{\nu}_L\gamma_\mu\nu_L)(\bar{\nu}_L\gamma^\mu\nu_L)$$

Need a cosmic  $\nu$  with E $_{\nu}$   $\sim$ 10 $^{11}$  GeV to scatter once off C $\nu$ B over distance H $^{-1}$  Weiler (1982)

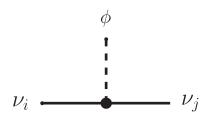
Lots of room for beyond SM contributions



### A wide range of motivations

### **Neutrino Mass Generation**

Models with a Higgs mechanism in the neutrino sector



See, e.g., Chikashige *et al* '81 Gu & He '07

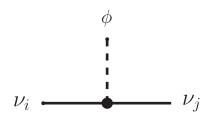
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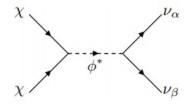
#### **Dark Matter**

Models with a Higgs mechanism in the neutrino sector

Predictive models of thermal (WIMP-like) or non-thermal (freeze-in) dark matter



See, e.g., Chikashige *et al* '81 Gu & He '07



See, e.g., Kelly & Zhang '19 de Gouvêa *et al* '19 Kelly *et al* '20

## A wide range of motivations

#### **Neutrino Mass Generation**

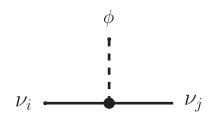
#### Dark Matter

#### **Observational Anomalies**

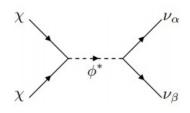
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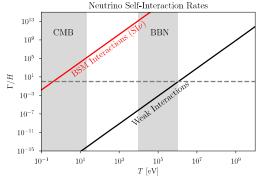
Hubble tension alleviated via neutrino self-interactions



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Kreisch, Cyr-Racine & Doré (2019)

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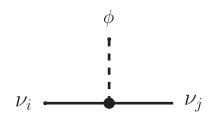
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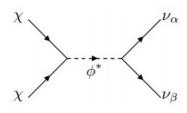
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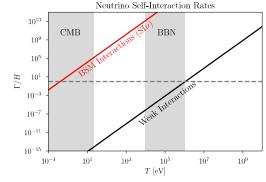
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### **Theoretical:**

U(1) anomaly-free extensions of SM include new neutrino couplings: B-L,  $L_i$ - $L_i$ ...



# **Observable Consequences**

### **Accelerators & Lab:**

- Rare decays of mesons and taus
- Oscillation and neutrino scattering exp.
- $0\nu\beta\beta$  searches
- ...

### **Astrophysics**

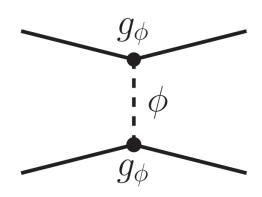
- Cosmic neutrino propagation
- Supernova physics
- Stellar evolution
- ...

### Cosmology:

- Microwave background
- Light element abundances
- DM clustering
- ...

# These span a huge range of energies, an EFT treatment is not possible

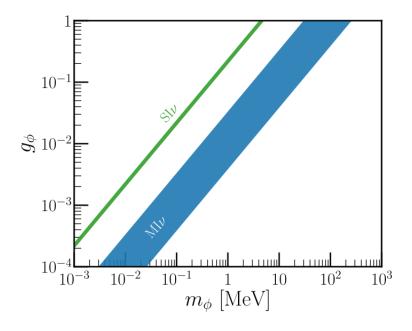
# **Complementarity Example**



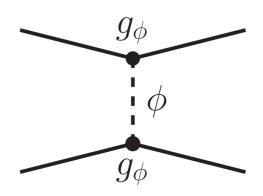
Hubble tension can be alleviated by neutrino selfinteractions during the CMB era

$$G_{\rm eff}({\rm SI}\nu) \sim 10^9 G_F$$

Kreisch, Cyr-Racine & Doré (2019)



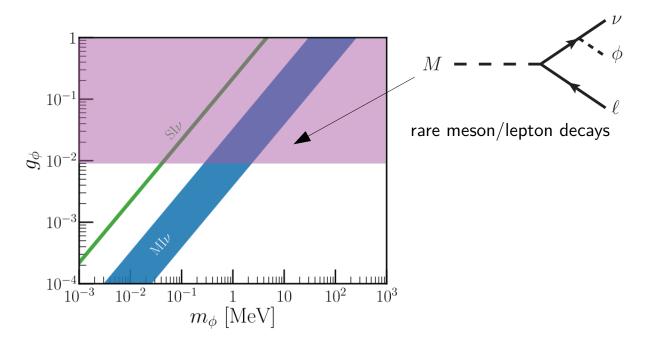
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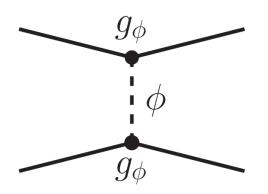
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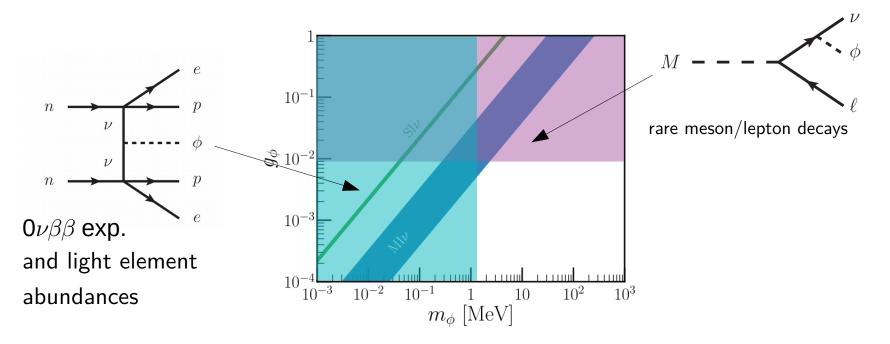
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# Goals for this Whitepaper

Define a set of benchmark models

 Comprehensively review phenomena within these common theoretical frameworks

Identify gaps in sensitivity and opportunities

Thank you!